

A Parametric Study of the SASI Comparing General Relativistic and Non-Relativistic Treatments

Samuel J. Dunham

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AND KELLY HOLLEY-BOCKELMANN ^{1,5}

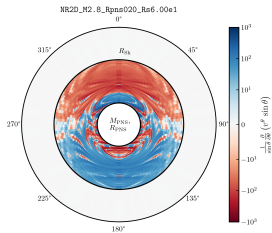
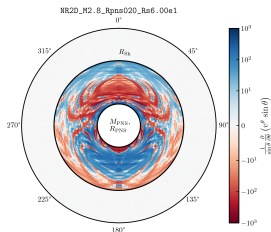
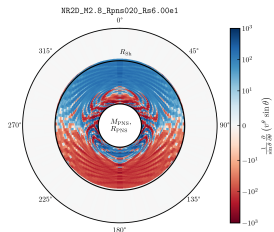
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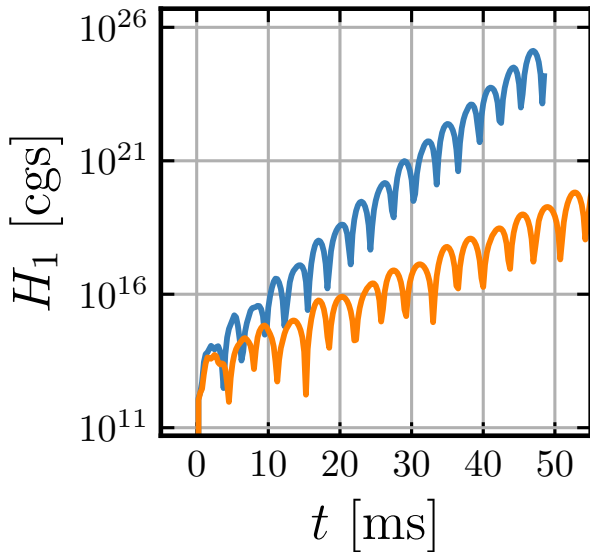
³*Computer Science and Mathematics Division, Oak Ridge National Laboratory, Oak Ridge, TN 37831*

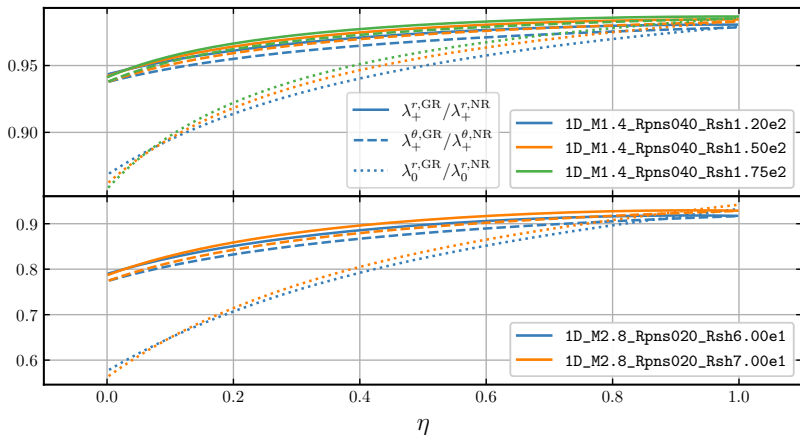
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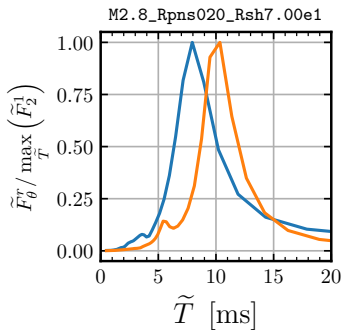
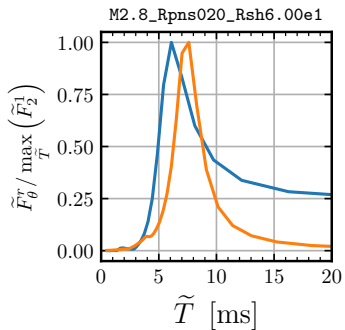
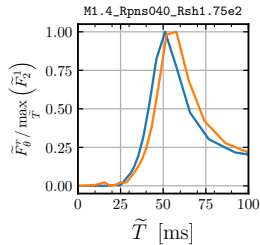
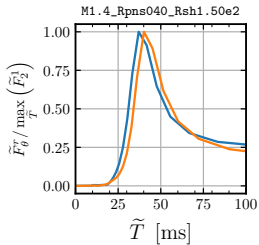
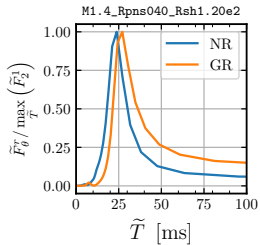
⁵*Department of Life and Physical Sciences, Fisk University, 1000 17th Ave N, Nashville, TN 37208*

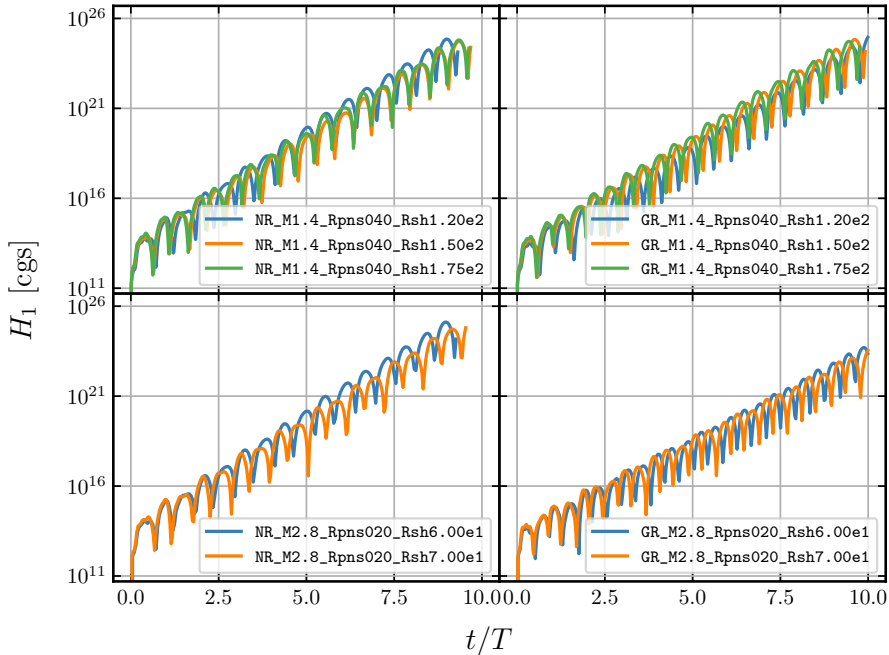
t  $t + 0.25 T$  $t + 0.5 T$ 

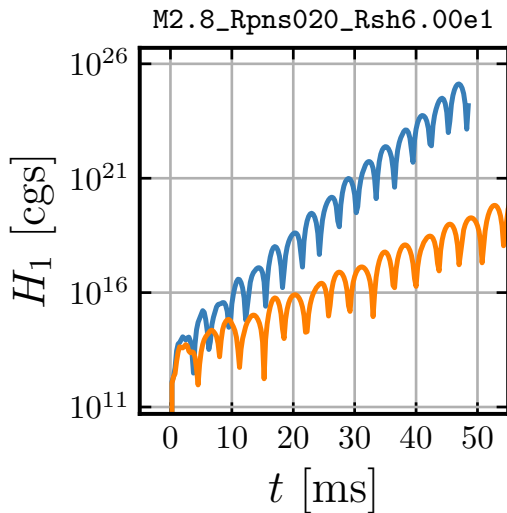
M2.8_Rpns020_Rsh6.00e1



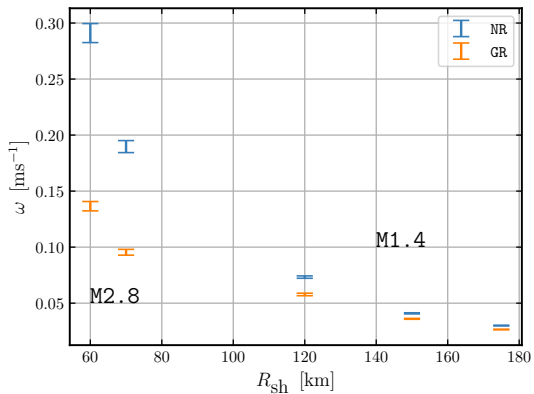






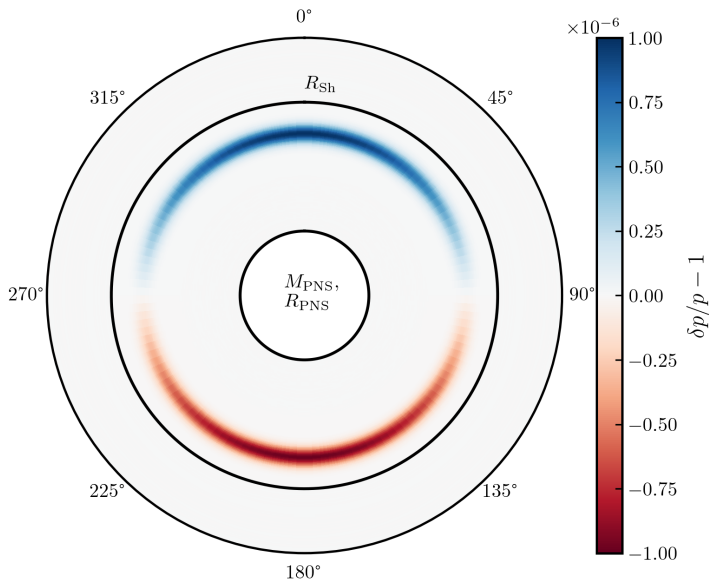


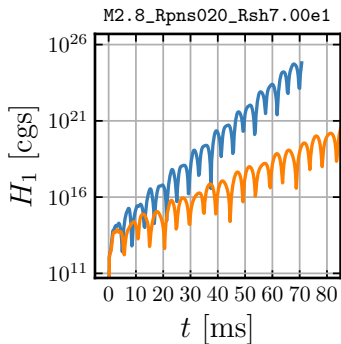
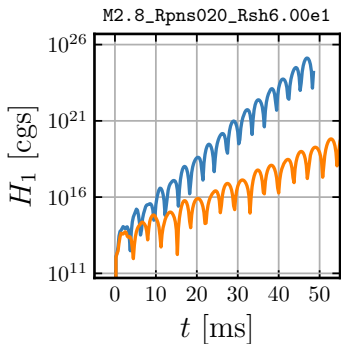
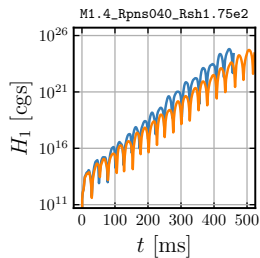
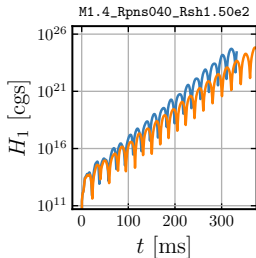
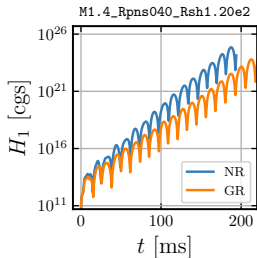
Conclusions



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- L. Scheck, H. Th. Janka, T. Foglizzo, and K. Kifonidis. Multidimensional supernova simulations with approximative neutrino transport. II. Convection and the advective-acoustic cycle in the supernova core. *A&A*, 477(3):931–952, January 2008. doi: 10.1051/0004-6361:20077701.
- Evan O'Connor and Christian D. Ott. Black Hole Formation in Failing Core-Collapse Supernovae. *ApJ*, 730(2):70, April 2011. doi: 10.1088/0004-637X/730/2/70.
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- Isabel Cordero-Carrión, Pablo Cerdá-Durán, Harald Dimmelmeier, José Luis Jaramillo, Jérôme Novak, and Ericourgoulhon. Improved constrained scheme for the Einstein equations: An approach to the uniqueness issue. *Phys. Rev. D*, 79(2):024017, January 2009. doi: 10.1103/PhysRevD.79.024017.
- Thomas W. Baumgarte and Stuart L. Shapiro. *Numerical Relativity: Solving Einstein's Equations on the Computer*. Cambridge, 2010.
- John M. Blondin, Anthony Mezzacappa, and Christine DeMarino. Stability of Standing Accretion Shocks, with an Eye toward Core-Collapse Supernovae. *ApJ*, 584(2):971–980, February 2003. doi: 10.1086/345812.

- Extended study of Blondin and Mezzacappa (2006) to include GR
- Showed that GR leads to longer SASI oscillation period than NR
- Showed that GR leads to smaller SASI growth rate than NR
- Found that growth rate is such that ωT is roughly constant for some parameter sets: implications for growth rate mechanism
- Future Work
 - Extend study to 3D
 - Include GR monopole (Marek et al., 2006)



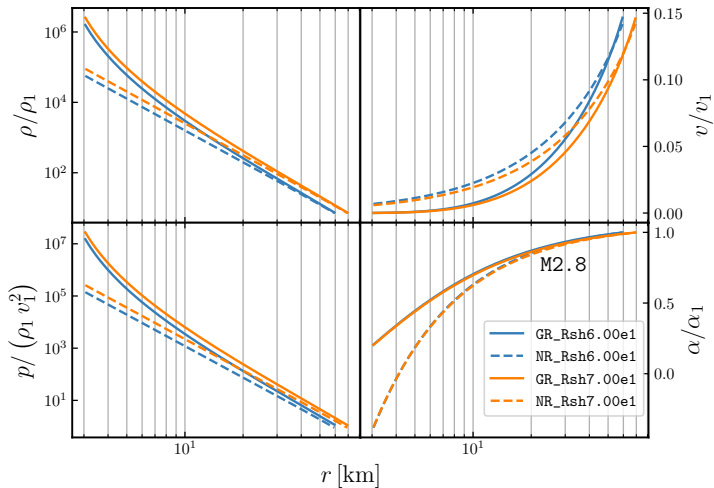


$$A := \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} (v^\theta \sin \theta) \quad (\text{Scheck et al., 2008})$$

$$A(r, \theta, t) = \sum_{\ell'=0}^{\infty} G_{\ell'}(r, t) P_{\ell'}(\cos \theta)$$

$$\implies G_\ell(r, t) := \frac{1}{N_\ell} \int_0^\pi A(r, \theta, t) P_\ell(\cos \theta) \sin \theta d\theta$$

$$H_\ell(t) := 4\pi \int_{r_a}^{r_b} [G_\ell(r, t)]^2 [\psi(r)]^6 r^2 dr$$



Parameters we varied:

$\xi =$

$(M_{\text{PNS}}/M_{\odot}) / (R_{\text{PNS}}/20 \text{ km})$

(O'Connor and Ott,

2011), $R_{\text{Sh}}(t = 0)$

Table 1. Model Parameters

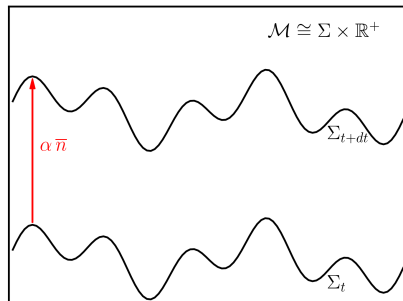
| Model | $M_{\text{PNS}} [M_{\odot}]$ | $R_{\text{PNS}} [\text{km}]$ | $R_{\text{sh}} [\text{km}]$ | ξ |
|------------------------|------------------------------|------------------------------|-----------------------------|-------|
| M1.4.Rpns040.Rsh1.20e2 | 1.4 | 40 | 120 | 0.7 |
| M1.4.Rpns040.Rsh1.50e2 | 1.4 | 40 | 150 | 0.7 |
| M1.4.Rpns040.Rsh1.75e2 | 1.4 | 40 | 175 | 0.7 |
| M2.8.Rpns020.Rsh6.00e1 | 2.8 | 20 | 60 | 2.8 |
| M2.8.Rpns020.Rsh7.00e1 | 2.8 | 20 | 70 | 2.8 |

NOTE—Model parameters chosen for the 5 models. All models were run with both GR and NR. The first three rows correspond to the low-compactness models and the last two rows correspond to the high-compactness models.

Conformally-Flat Condition

Wilson et al. (1996); Cordero-Carrión et al. (2009)

$$ds^2 = -\alpha^2 dt^2 + \gamma_{ij} dx^i dx^j$$



Conformally-Flat Condition

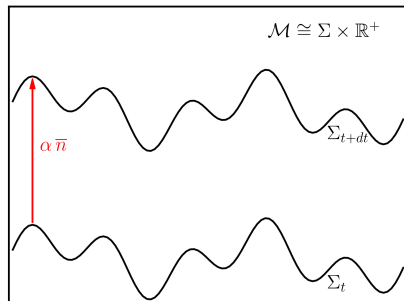
Wilson et al. (1996); Cordero-Carrión et al. (2009)

$$ds^2 = -\alpha^2 dt^2 + \gamma_{ij} dx^i dx^j$$

$$\gamma_{ij} = \psi^4 \bar{\gamma}_{ij}$$

$1 \leq \psi < 2$: Conformal factor

$$\bar{\gamma}_{ij} = \text{diag}(1, r^2, r^2 \sin^2 \theta)$$



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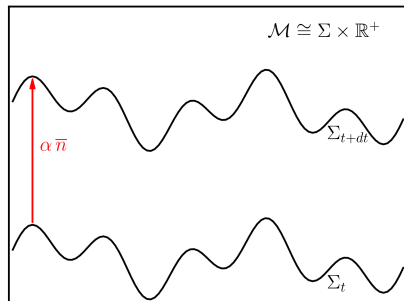
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(Also, maximum slicing condition:

$$K := \text{Tr}_{\gamma_{ij}} (\underline{\underline{K}}) = \partial_t K = 0)$$



Baumgarte and Shapiro (2010)

$$\alpha(r) = \left(1 - \frac{R_{\text{Sc}}}{r}\right) \left(1 + \frac{R_{\text{Sc}}}{r}\right)^{-1}$$

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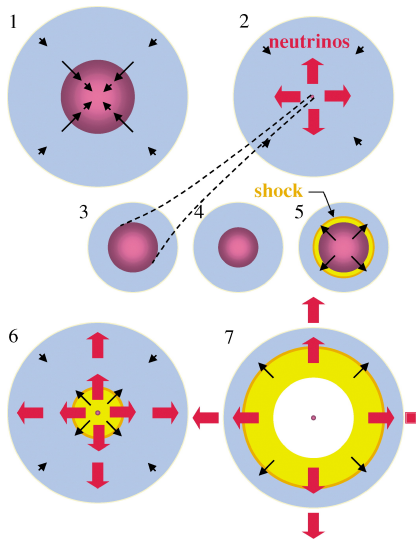
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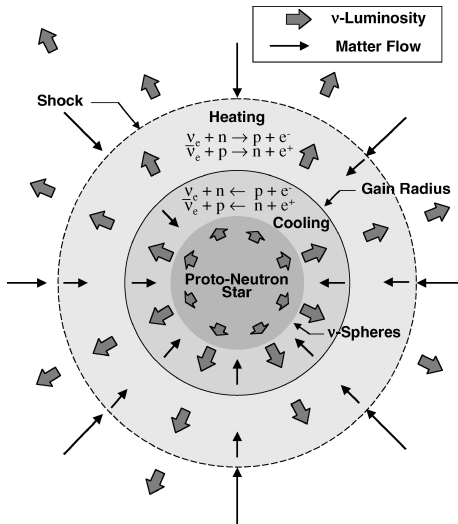
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$$\beta^i = 0$$

$$K_{ij} = 0$$



Mezzacappa A. 2005.
 Annu. Rev. Nucl. Part. Sci. 55:467–515



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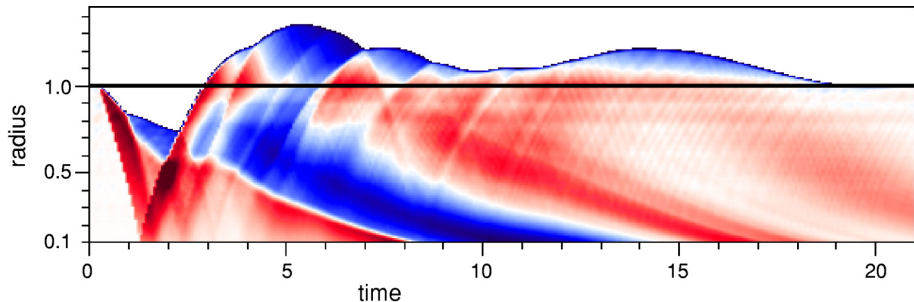


Figure: Equilibrium (white), under- (blue), and over- (red) pressure (Blondin et al., 2003).

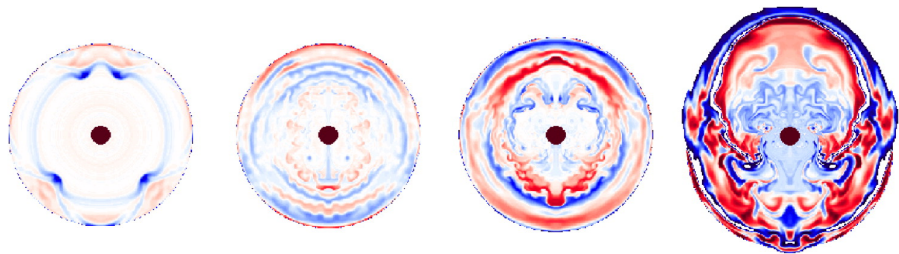
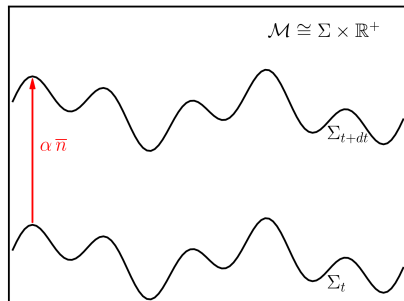


Figure: Equilibrium (white), under- (blue), and over- (red) entropy (Blondin et al., 2003).

$d+1$ Decomposition

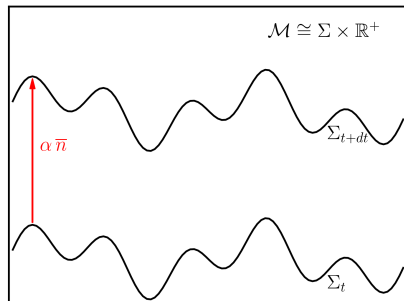
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g : spacetime metric on \mathcal{M}

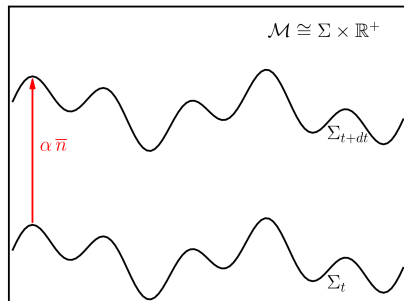


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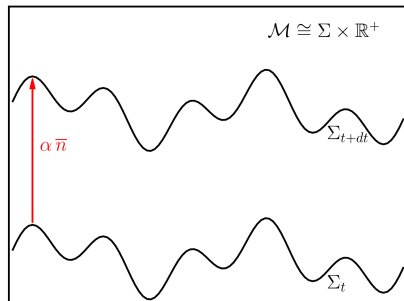
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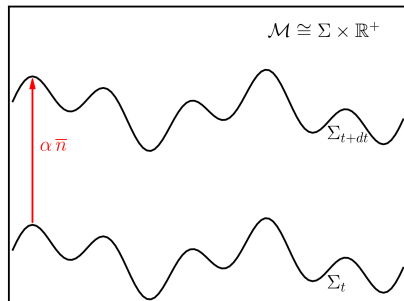
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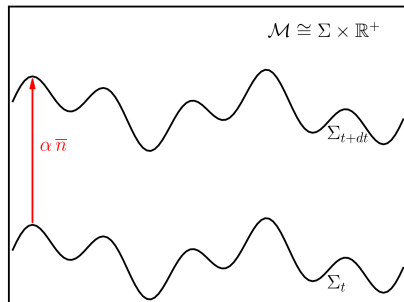
$$g_{\mu\nu} = \begin{pmatrix} -\alpha^2 + \beta^k \beta_k & \beta_j \\ \beta_i & \gamma_{ij} \end{pmatrix}$$

$\underline{\gamma}$: spatial three-metric on Σ_t

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\bar{n} : Eulerian four-velocity

$$(\underline{g}(\bar{n}, \bar{n}) = \bar{n} \cdot \bar{n} = -1)$$



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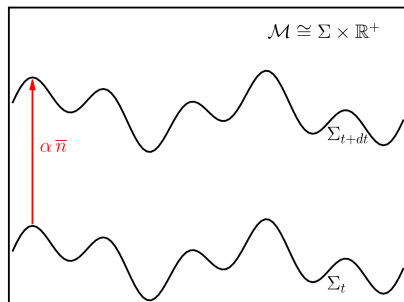
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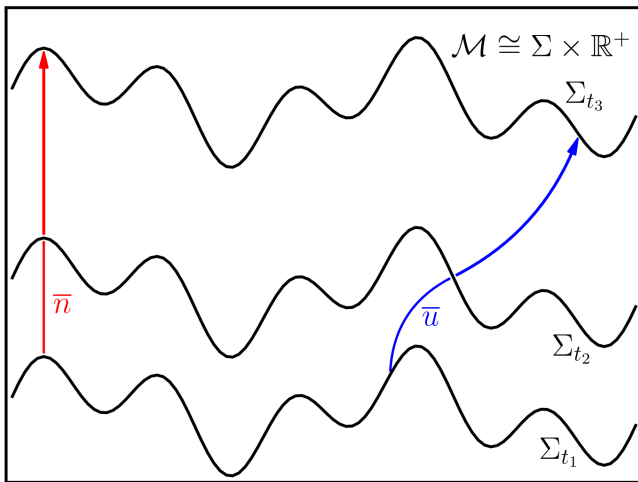
$0 < \alpha \leq 1$: Lapse Function

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$$(\underline{g}(\bar{n}, \bar{n}) = \bar{n} \cdot \bar{n} = -1)$$

(Also, \underline{K} : Extrinsic curvature)





Fluid Equations

Units defined such that $c = G = 1$

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$$\bar{\nabla} \cdot \bar{J} = 0 \quad (\bar{J}: \text{baryon mass density current four-vector})$$

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$$\bar{\bar{T}} := \rho h \bar{u} \otimes \bar{u} + p \bar{\bar{g}} \quad (p: \text{comoving thermal pressure},$$

$$h := 1 + (e + p) / \rho: \text{specific enthalpy, } e: \text{comoving internal energy density})$$

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Five equations with six unknowns ☹

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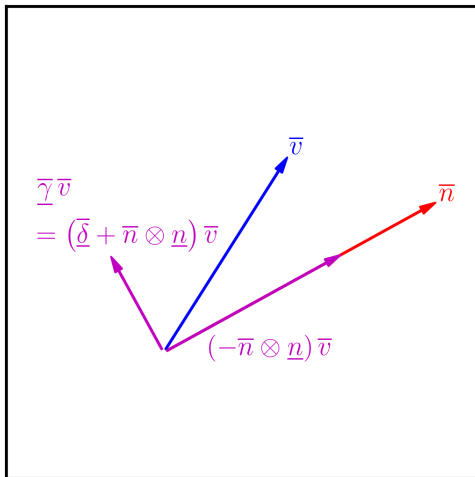
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Five equations with six unknowns ☹

Close with an equation of state: $p = p(e) := (\Gamma - 1)e$, $\Gamma = 4/3$

Valencia Decomposition



Extensible to higher-rank tensors!

$$E := n_{\mu'} n_{\nu'} T^{\mu'\nu'}$$

Valencia Decomposition

$$E := n_{\mu'} n_{\nu'} T^{\mu'\nu'}$$

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Valencia Decomposition

$$E := n_{\mu'} n_{\nu'} T^{\mu'\nu'}$$

$$S^{\mu} := -\gamma^{\mu}_{\mu'} n_{\nu'} T^{\mu'\nu'}$$

$$P^{\mu\nu} := \gamma^{\mu}_{\mu'} \gamma^{\nu}_{\nu'} T^{\mu'\nu'}$$

Math...

$$\partial_t \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_i [\alpha \sqrt{\gamma} \mathbf{F}^i(\mathbf{U})] = \mathbf{S}(\mathbf{U})$$

GR

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GR

$$\mathbf{U} := (D, S_j, \tau)^\top$$

$$\partial_t \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_i [\alpha \sqrt{\gamma} \mathbf{F}^i(\mathbf{U})] = \mathbf{S}(\mathbf{U})$$

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$$\mathbf{U} := (D, S_j, \tau)^\top$$

$$D := \rho W$$

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GR

$$\mathbf{U} := (D, S_j, \tau)^\top$$

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$$\partial_t \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_i [\alpha \sqrt{\gamma} \mathbf{F}^i(\mathbf{U})] = \mathbf{S}(\mathbf{U})$$

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$$\tau := E - D = \rho h W^2 - p - \rho W$$

$$\partial_t \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_i [\alpha \sqrt{\gamma} \mathbf{F}^i(\mathbf{U})] = \mathbf{S}(\mathbf{U})$$

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$$\alpha = (1 - R_{\text{Sc}}/r) / (1 + R_{\text{Sc}}/r)$$

$$\sqrt{\gamma} = \psi^6 \sqrt{\bar{\gamma}}$$

$$\partial_t \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_i [\alpha \sqrt{\gamma} \mathbf{F}^i(\mathbf{U})] = \mathbf{S}(\mathbf{U})$$

GR

NR

$$\mathbf{U} := (D, S_j, \tau)^\top$$

$$D := \rho W$$

$$S_j := \rho h W^2 v_j$$

$$\tau := E - D = \rho h W^2 - p - \rho W$$

$$\alpha = (1 - R_{\text{Sc}}/r) / (1 + R_{\text{Sc}}/r)$$

$$\sqrt{\gamma} = \psi^6 \sqrt{\bar{\gamma}}$$

$$\partial_t \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_i [\alpha \sqrt{\gamma} \mathbf{F}^i(\mathbf{U})] = \mathbf{S}(\mathbf{U})$$

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$$\tau = e + \frac{1}{2} \rho v^i v_i$$

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$$S_j := \rho v_j$$

$$\tau = e + \frac{1}{2} \rho v^i v_i$$

$$\alpha = 1$$

$$\partial_t \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_i [\alpha \sqrt{\gamma} \mathbf{F}^i(\mathbf{U})] = \mathbf{S}(\mathbf{U})$$

GR

$$\mathbf{U} := (D, S_j, \tau)^\top$$

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$$\partial_t \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_i [\alpha \sqrt{\gamma} \mathbf{F}^i(\mathbf{U})] = \mathbf{S}(\mathbf{U})$$

GR

$$\mathbf{F}^i(\mathbf{U}) = \begin{pmatrix} \rho W v^i \\ \rho h W^2 v^i v_j + p \delta^i_j \\ (\rho h W^2 - \rho W) v^i \end{pmatrix}$$

NR

$$\mathbf{F}^i(\mathbf{U}) = \begin{pmatrix} \rho v^i \\ \rho v^i v_j + p \delta^i_j \\ (\rho h_{\text{NR}} + \frac{1}{2} v^j v_j) v^i \end{pmatrix}$$

$$\partial_t \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_i [\alpha \sqrt{\gamma} \mathbf{F}^i(\mathbf{U})] = \mathbf{S}(\mathbf{U})$$

GR

$$\mathbf{S}(\mathbf{U}) = \begin{pmatrix} 0 \\ \frac{1}{2} \alpha P^{ik} \partial_j \gamma_{ik} - E \partial_j \alpha \\ -S^j \partial_j \alpha \end{pmatrix}$$

NR

$$\mathbf{S}(\mathbf{U}) = \begin{pmatrix} 0 \\ \frac{1}{2} P^{ik} \partial_j \bar{\gamma}_{ik} - \rho \partial_j \Phi \\ -S^j \partial_j \Phi \end{pmatrix}$$
$$\Phi(r) := -M/r$$

Steady-State Solutions (pre-shock)

$$\partial_t \mathbf{U}^0 + \frac{1}{\sqrt{\gamma}} \partial_r [\alpha \sqrt{\gamma} \mathbf{F}^r(\mathbf{U})] = \mathbf{S}(\mathbf{U})$$

GR

Steady-State Solutions (pre-shock)

$$\partial_t \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_r [\alpha \sqrt{\gamma} \mathbf{F}^r(\mathbf{U})] = \mathbf{S}(\mathbf{U})$$

GR

$$\alpha \psi^6 W \times 4\pi r^2 \rho v = -\dot{M}$$

Steady-State Solutions (pre-shock)

$$\partial_t \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_r [\alpha \sqrt{\gamma} \mathbf{F}^r(\mathbf{U})] = \mathbf{S}(\mathbf{U})$$

GR

$$\alpha \psi^6 W \times 4\pi r^2 \rho v = -\dot{M}$$

$$\alpha h W = \mathcal{B}$$

Steady-State Solutions (pre-shock)

$$\partial_t \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_r [\alpha \sqrt{\gamma} \mathbf{F}^r(\mathbf{U})] = \mathbf{S}(\mathbf{U})$$

GR

$$\alpha \psi^6 W \times 4\pi r^2 \rho v = -\dot{M}$$

$$\alpha h W = \mathcal{B}$$

$$p = K_1 \rho^\Gamma$$

Steady-State Solutions (pre-shock)

$$\partial_t \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_r [\alpha \sqrt{\gamma} \mathbf{F}^r(\mathbf{U})] = \mathbf{S}(\mathbf{U})$$

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$$\partial_t \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_r [\alpha \sqrt{\gamma} \mathbf{F}^r(\mathbf{U})] = \mathbf{S}(\mathbf{U})$$

GR

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$$\alpha h W = \mathcal{B}$$

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NR

$$4\pi r^2 \rho v = -\dot{M}$$

Steady-State Solutions (pre-shock)

$$\partial_t \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_r [\alpha \sqrt{\gamma} \mathbf{F}^r(\mathbf{U})] = \mathbf{S}(\mathbf{U})$$

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$$\alpha h W = \mathcal{B}$$

$$p = K_1 \rho^\Gamma$$

NR

$$4\pi r^2 \rho v = -\dot{M}$$

$$\frac{1}{2} v^2 + h_{\text{NR}} + \Phi = \mathcal{B}_{\text{NR}}$$

Steady-State Solutions (pre-shock)

$$\partial_t \mathbf{U}^0 + \frac{1}{\sqrt{\gamma}} \partial_r [\alpha \sqrt{\gamma} \mathbf{F}^r(\mathbf{U})] = \mathbf{S}(\mathbf{U})$$

GR

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$$\alpha h W = \mathcal{B}$$

$$p = K_1 \rho^\Gamma$$

NR

$$4\pi r^2 \rho v = -\dot{M}$$

$$\frac{1}{2} v^2 + h_{\text{NR}} + \Phi = \mathcal{B}_{\text{NR}}$$

$$p = K_1 \rho^\Gamma$$

Jump Conditions

$$U_1 \neq U_2$$

$$F^r(U_1) = F^r(U_2)$$

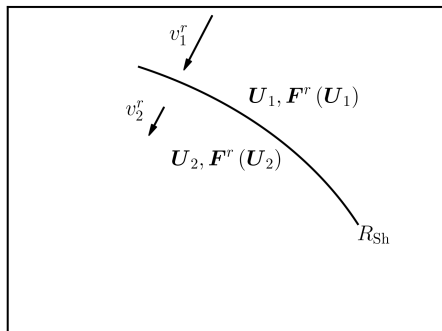
Yields:

$$\rho_2$$

$$e_2(p_2)$$

$$K_2 (> K_1)$$

$$v_2^r$$



Steady-State Solutions (post-shock)

$$\partial_t \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_r [\alpha \sqrt{\gamma} \mathbf{F}^r(\mathbf{U})] = \mathbf{S}(\mathbf{U})$$

GR

$$\alpha \psi^6 W \times 4\pi r^2 \rho v = -\dot{M}$$

$$\alpha h W = \mathcal{B}$$

$$p = K_2 \rho^\Gamma$$

NR

$$4\pi r^2 \rho v = -\dot{M}$$

$$\frac{1}{2} v^2 + h_{\text{NR}} + \Phi = \mathcal{B}_{\text{NR}}$$

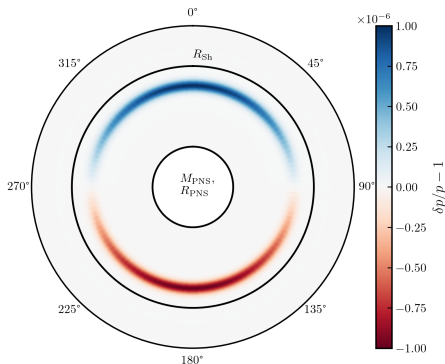
$$p = K_2 \rho^\Gamma$$

$$\eta(r) := \frac{r - R_{\text{PNs}}}{R_{\text{Sh}} - R_{\text{PNs}}}$$

$$\frac{\delta p(\eta, \theta)}{p(\eta_c)} = 10^{-6} \times \exp\left[\frac{-(\eta - \eta_c)^2}{2\sigma^2}\right] \cos \theta$$

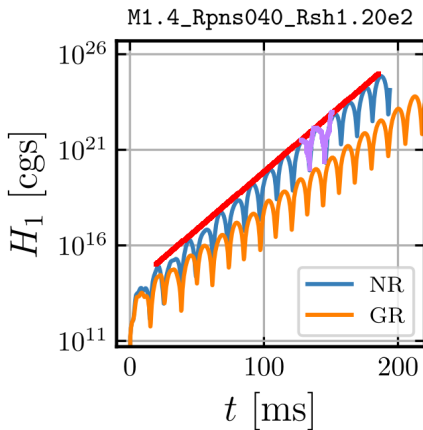
$$\eta_c = 0.75$$

$$\sigma = 0.05$$



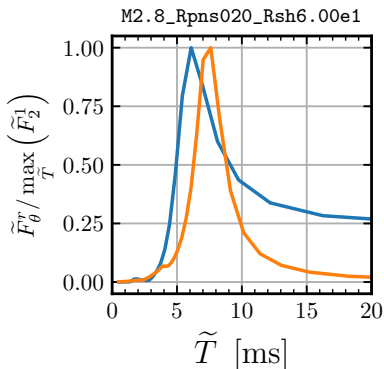
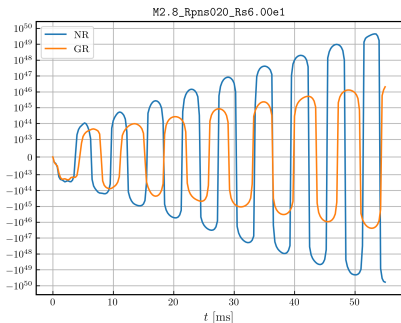
$$F(t) = F(0) e^{2\omega t} \sin^2\left(\frac{2\pi t}{T} + \delta\right)$$

(Blondin and Mezzacappa, 2006)



$$F_\theta^r := \alpha \psi^6 h W^2 \times \sqrt{\gamma} \rho v^r v_\theta$$

$$\tilde{F}_\theta^r := \text{FFT} \{F_\theta^r\}$$



T defined as the unique \tilde{T} such that $\tilde{F}_\theta^r(\tilde{T}) = 1$

